

ϕ -CONHARMONICALLY SYMMETRIC SASAKIAN MANIFOLDS

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ABSTRACT. We consider some conditions on conharmonic curvature tensor \tilde{C} , which has many applications in physics and mathematics. We prove that every ϕ -conharmonically symmetric n -dimensional ($n > 3$), Sasakian manifold is an Einstein manifold. Also we prove that a three-dimensional Sasakian manifold is locally ϕ -conharmonically symmetric if and only if it is locally ϕ -symmetric. Finally we give two examples of a three-dimensional ϕ -conharmonically symmetric Sasakian manifold.

1. INTRODUCTION

Let (M^n, g) be an n -dimensional, $n \geq 3$, Riemannian manifold of class C^∞ . The conharmonic curvature tensor \tilde{C} is considered as an invariant of the conharmonic transformation defined by Ishii [6]. It satisfies all the symmetric properties of the Riemannian curvature tensor R . There are many physical applications of the tensor \tilde{C} . For example, in [1], Abdussattar showed that sufficient condition for a space-time to be conharmonic to a flat space-time is that the tensor \tilde{C} vanishes identically. A conharmonically flat space-time is either empty in which case it is flat or is filled with a distribution represented by energy momentum tensor T possessing the algebraic structure of an electromagnetic field and is conformal to flat space-time [1]. Also he described the gravitational field due to a distribution of pure radiation in presence of disordered radiation by means of spherically symmetric conharmonically flat space-time.

On the other hand, the notion of locally symmetry of a Riemannian manifold has been weakened by many authors in several ways to different extent. As a weaker version of locally symmetry, T. Takashi [7] introduced the notion of locally ϕ -symmetry on a Sasakian manifold. In the context of contact geometry the notion of ϕ -symmetry is introduced and studied by Boeckx, Buecken and Vanhacck [5]. In [4], Boeckx proved that every non-Sasakian (κ, μ) - manifold is locally ϕ -symmetric in the strong sense.

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In the present work we study ϕ -conharmonically symmetry in a Sasakian manifold. The paper is organized as follows: In Section 2, we give a brief account of conharmonic curvature tensor, Weyl tensor and Sasakian manifold. In Section 3, we consider ϕ -conharmonically symmetric Sasakian manifold and prove that it is an Einstein manifold. Then using this result we concluded that a Sasakian manifold is ϕ -conharmonically symmetric if and only if it is ϕ -symmetric. In the next section we consider three-dimensional locally ϕ -conharmonically symmetric Sasakian manifold. Finally we give two examples of a three-dimensional ϕ -conharmonically symmetric Sasakian manifold.

2. Preliminaries

In this section, we collect some basic facts about contact metric manifolds. We refer to [3] for a more detailed treatment. An n -dimensional ($n = 2m + 1$) differentiable manifold M^n is called a *contact manifold* if there exists a globally defined 1-form η such that $(d\eta)^m \wedge \eta \neq 0$. On a contact manifold there exists a unique global vector field ξ satisfying

$$d\eta(\xi, X) = 0, \quad \eta(\xi) = 1, \quad (2.1)$$

for any vector field X tangent to M .

Moreover, it is well-known that there exist a $(1, 1)$ -tensor field ϕ , a Riemannian metric g which satisfy

$$\phi^2 = -I + \eta \otimes \xi, \quad (2.2)$$

$$g(\phi X, \phi Y) = g(X, Y) - \eta(X)\eta(Y), \quad g(\xi, X) = \eta(X), \quad (2.3)$$

$$d\eta(X, Y) = g(X, \phi Y), \quad (2.4)$$

for all X, Y tangent to M . As a consequence of the above relations we have

$$\phi\xi = 0, \quad \eta\phi = 0. \quad (2.5)$$

The structure (ϕ, ξ, η, g) is called a *contact metric structure* and the manifold M^n with a contact metric structure is said to be a *contact metric manifold*. Furthermore, if moreover the structure is normal, that is, $[\phi X, \phi Y] + \varphi^2[X, Y] - \phi[X, \phi Y] - \phi[\phi X, Y] = -2d\eta(X, Y)\xi$, then the contact metric structure is called a *Sasakian structure* (normal contact metric structure) and M is called a *Sasakian manifold*.

We denote by ∇ the Levi-Civita connection on M . Then we have

$$(i)(\nabla_X \phi)Y = g(X, Y)\xi - \eta(Y)X, \quad (ii)\nabla_X \xi = -\phi X, \quad (2.6)$$

$$(i)R(X, Y)\xi = \eta(Y)X - \eta(X)Y, \quad (ii)S(X, \xi) = 2n\eta(X) \quad (2.7)$$

for any vector fields X, Y tangent to M , where S denotes the Ricci tensor [3].

The *Weyl conformal curvature tensor* C and the *conharmonic curvature tensor* \tilde{C} are defined by

$$\begin{aligned} C(X, Y)Z &= R(X, Y)Z - \frac{1}{n-2} \left[\begin{array}{l} g(Y, Z)QX - g(X, Z)QY \\ + S(Y, Z)X - S(X, Z)Y \end{array} \right] \\ &\quad + \frac{r}{(n-1)(n-2)} [g(Y, Z)Z - g(X, Z)Y] \end{aligned} \quad (2.8)$$

and

$$\tilde{C}(X, Y)Z = R(X, Y)Z - \frac{1}{n-2} \left[\begin{array}{l} g(Y, Z)QX - g(X, Z)QY \\ + S(Y, Z)X - S(X, Z)Y \end{array} \right] \quad (2.9)$$

respectively, where Q denotes the Ricci operator, i.e. $S(X, Y) = g(QX, Y)$ and r is scalar curvature [6]. The curvature tensor R of a 3-dimensional Riemannian manifold can be written as

$$\begin{aligned} R(X, Y)Z &= g(Y, Z)QX - g(X, Z)QY + S(Y, Z)X - S(X, Z)Y \\ &\quad - \frac{r}{2}(g(Y, Z)X - g(X, Z)Y). \end{aligned} \quad (2.10)$$

3. ϕ -conharmonically symmetric Sasakian manifolds

Definition 3.1. A Sasakian manifold M^n is said to be ϕ -symmetric if

$$\phi^2(\nabla_X R)(Y, Z)W = 0,$$

for any vector fields X, Y, Z, W of M . If the vector fields are orthogonal to ξ , then the manifold is called locally ϕ -symmetric.

Definition 3.2. A Sasakian manifold $M^n(\phi, \xi, \eta, g)$ is said to be ϕ -conharmonically symmetric if

$$\phi^2(\nabla_X \tilde{C})(Y, Z)W = 0, \quad (3.1)$$

for any vector fields X, Y, Z, W of M . If the vector fields are orthogonal to ξ , then the manifold is called locally ϕ -conharmonically symmetric.

From the definition it follows that a ϕ -symmetric Sasakian manifold is ϕ -conharmonically symmetric. But the converse is not true in general.

Firstly, differentiating (2.9) covariantly with respect to X , we obtain

$$(\nabla_X \tilde{C})(Y, Z)W = (\nabla_X R)(Y, Z)W \quad (3.2)$$

$$- \frac{1}{n-2} [(\nabla_X S)(Z, W)Y - (\nabla_X S)(Y, W)Z + g(Z, W)(\nabla_X Q)Y - g(Y, W)(\nabla_X Q)Z].$$

Using (3.1) and (2.2), we get

$$\begin{aligned} & -g((\nabla_X R)(Y, Z)W, U) + \frac{1}{n-2} [(\nabla_X S)(Z, W)g(Y, U) - (\nabla_X S)(Y, W)g(Z, U) \\ & + g(Z, W)g((\nabla_X Q)Y, U) - g(Y, W)g((\nabla_X Q)Z, U)] + g((\nabla_X R)(Y, Z)W, \xi)\xi \\ & + \frac{1}{n-2} [g((\nabla_X S)(Z, W)Y - (\nabla_X S)(Y, W)Z; \xi)\eta(U) \\ & + g(Z, W)g((\nabla_X Q)Y, \xi)\eta(U) - g(Y, W)g((\nabla_X Q)Z, \xi)\eta(U)] = 0. \end{aligned} \quad (3.3)$$

Applying contraction to the equation (3.3) with respect to Y and U , we have

$$\begin{aligned} & -(\nabla_X S)(Z, W) + \frac{1}{n-2} [(n-2)(\nabla_X S)(Z, W) + g(Z, W)X(r)] \\ & + g((\nabla_X R)(\xi, Z)W, \xi) - \frac{1}{n-2} [(\nabla_X S)(Z, W) - (\nabla_X S)(\xi, W)\eta(Z) \\ & + g(Z, W)g((\nabla_X Q)\xi, \xi) - \eta(W)g((\nabla_X Q)Z, \xi)] = 0. \end{aligned} \quad (3.4)$$

Taking $W = \xi$ in (3.4) it follows that

$$-(\nabla_X S)(Z, \xi) + \frac{1}{n-2}\eta(Z)X(r) = 0. \quad (3.5)$$

Then putting $Z = \xi$ in (3.5), we obtain $X(r) = 0$, that is, r is constant. Thus we can state the following:

Theorem 1. Let M be a Sasakian manifold. If M is ϕ -conharmonically symmetric then the scalar curvature r is constant.

From the equation (3.5) and Theorem 1 we obtain

$$(\nabla_X S)(Z, \xi) = 0,$$

that is,

$$\nabla_X S(Z, \xi) - S(\nabla_X Z, \xi) - S(Z, \nabla_X \xi) = 0.$$

Now using 6(ii) and 7(ii) yields

$$2n(\nabla_X \eta)(Z) + S(Z, \phi X) = 0. \quad (3.6)$$

Also in a Sasakian manifold it is known that $(\nabla_X \eta)(Z) = g(X, \phi Z)$. Therefore putting $X = \phi X$ in (3.6) we get

$$S(X, Z) = 2ng(X, Z).$$

Hence we are in a position to state the following:

Theorem 2. *Let M be a Sasakian manifold. If M is ϕ -conharmonically symmetric then M is an Einstein manifold.*

Then using the above theorem in the equation (3.2), we get easily $(\nabla_X \tilde{C})(Y, Z)W = (\nabla_X R)(Y, Z)W$. So, we state the following:

Corollary 1. *Let M^n be a Sasakian manifold. M^n is ϕ -conharmonically symmetric if and only if it is ϕ -symmetric.*

4. Three-dimensional locally ϕ -conharmonically symmetric Sasakian manifolds

Now, we suppose that M is a three-dimensional locally ϕ -conharmonically symmetric Sasakian manifold. Using the equation (2.9), we get

$$\phi^2(\nabla_X \tilde{C})(Y, Z)W = -\frac{X(r)}{2}[g(Y, W)Z - g(Z, W)Y]$$

for any vector fields X, Y, Z, W are orthogonal to ξ . Thus we can easily get the following:

Theorem 3. *A three-dimensional Sasakian manifold is locally ϕ -conharmonically symmetric if and only if the scalar curvature r is constant.*

It is known from Watanabe's result [9] that a three-dimensional Sasakian manifold is locally ϕ -symmetric if and only if the scalar curvature r is constant. Using Watanabe's result we state the following:

Theorem 4. *A three-dimensional Sasakian manifold is locally ϕ -conharmonically symmetric if and only if it is locally ϕ -symmetric.*

5. Example

In this section we give two examples to prove the existence of a three-dimensional ϕ -conharmonically symmetric Sasakian manifold.

Example 5.1. *In [8](p.275), K.Yano and M.Kon gave an example of a Sasakian manifold which is three-dimensional sphere. Three-dimensional sphere is an Einstein manifold and hence a manifold of constant scalar curvature. Hence by Theorem 3 the three-dimensional sphere is locally ϕ -conharmonically symmetric.*

Example 5.2. *We consider the three-dimensional manifold $M = \{(x, y, z) \in \mathbb{R}^3, (x, y, z) \neq 0\}$, where (x, y, z) are standard coordinate of \mathbb{R}^3 .*

The vector fields

$$e_1 = \frac{\partial}{\partial x} - y \frac{\partial}{\partial z}, \quad e_2 = \frac{\partial}{\partial y}, \quad e_3 = \frac{1}{2} \frac{\partial}{\partial z}$$

are linearly independent at each point of M .

Let g be the Riemannian metric defined by

$$g(e_1, e_3) = g(e_1, e_2) = g(e_2, e_3) = 0,$$

$$g(e_1, e_1) = g(e_2, e_2) = g(e_3, e_3) = 1,$$

that is, the form of the metric becomes

$$g = \frac{dx^2 + dy^2 + dz^2}{z^2}.$$

Let η be the 1-form defined by $\eta(Z) = g(Z, e_3)$ for any $Z \in \chi(M)$.

Further, let ϕ be the $(1, 1)$ tensor field defined by

$$\phi(e_1) = e_2, \quad \phi(e_2) = -e_1, \quad \phi(e_3) = 0.$$

So, using the linearity of ϕ and g , we have

$$\eta(e_3) = 1,$$

$$\phi^2 Z = -Z + \eta(Z)e_3,$$

$$g(\phi Z, \phi W) = g(Z, W) - \eta(Z)\eta(W),$$

for any $Z, W \in \chi(M)$.

Then for $e_3 = \xi$, the structure (ϕ, ξ, η, g) defines an almost contact metric structure on M .

Let ∇ be the Levi-Civita connection with respect to metric g . Then we have

$$\begin{aligned} [e_1, e_2] &= e_1 e_2 - e_2 e_1 \\ &= \left(\frac{\partial}{\partial x} - y \frac{\partial}{\partial z} \right) \frac{\partial}{\partial y} - \frac{\partial}{\partial y} \left(\frac{\partial}{\partial x} - y \frac{\partial}{\partial z} \right) \\ &= \frac{\partial^2}{\partial x \partial y} - y \frac{\partial^2}{\partial z \partial y} - \frac{\partial^2}{\partial y \partial x} + \frac{\partial}{\partial z} + y \frac{\partial^2}{\partial y \partial z} \\ &= \frac{\partial}{\partial z} = 2e_3. \end{aligned}$$

Similarly

$$[e_2, e_3] = 0 \quad \text{and} \quad [e_1, e_3] = 0.$$

The Riemannian connection ∇ of the metric g is given by

$$\begin{aligned} 2g(\nabla_X Y, Z) &= Xg(Y, Z) + Yg(Z, X) - Zg(X, Y) \\ &\quad + g(Z, [X, Y]) - g(X, [Y, Z]) + g(Y, [Z, X]), \end{aligned} \quad (5.1)$$

which is known as Koszul's formula. Using (5.1) we have

$$2g(\nabla_{e_1} e_3, e_1) = 0 = 2g(-e_2, e_1). \quad (5.2)$$

Again by (5.1)

$$\begin{aligned} 2g(\nabla_{e_1} e_3, e_2) &= g(-2e_3, e_3) \\ &= 2g(-e_2, e_2) \end{aligned} \quad (5.3)$$

and

$$2g(\nabla_{e_1} e_3, e_3) = 0 = 2g(-e_2, e_3). \quad (5.4)$$

From (5.2), (5.3) and (5.4) we obtain

$$2g(\nabla_{e_1}e_3, X) = 2g(-e_2, X)$$

for all $X \in \chi(M)$.

Thus

$$\nabla_{e_1}e_3 = -e_2.$$

Therefore, (5.1) further yields

$$\begin{aligned} \nabla_{e_1}e_3 &= -e_2, & \nabla_{e_1}e_2 &= e_3, & \nabla_{e_1}e_1 &= 0 \\ \nabla_{e_2}e_3 &= e_1, & \nabla_{e_2}e_2 &= 0, & \nabla_{e_2}e_1 &= -e_3 \\ \nabla_{e_3}e_3 &= 0, & \nabla_{e_3}e_2 &= e_1, & \nabla_{e_3}e_1 &= -e_2 \end{aligned} \quad (5.5)$$

(5.5) tells us that the (ϕ, ξ, η, g) structure satisfies the formula $\nabla_X\xi = -\phi X$ for $\xi = e_3$. Hence $M(\phi, \xi, \eta, g)$ is a three-dimensional Sasakian manifold.

It is known that

$$R(X, Y)Z = \nabla_X\nabla_YZ - \nabla_Y\nabla_XZ - \nabla_{[X, Y]}Z. \quad (5.6)$$

With the help of the above results and using (5.6) it can be easily verified that

$$\begin{aligned} R(e_1, e_2)e_3 &= 0, & R(e_2, e_3)e_3 &= e_2, & R(e_1, e_3)e_3 &= e_1 \\ R(e_1, e_2)e_2 &= -3e_1, & R(e_2, e_3)e_2 &= -e_3, & R(e_1, e_3)e_2 &= 0 \\ R(e_1, e_2)e_1 &= 3e_2, & R(e_2, e_3)e_1 &= 0, & R(e_1, e_3)e_1 &= -e_3. \end{aligned}$$

From the above expression of the curvature tensor we obtain

$$S(e_1, e_1) = g(R(e_1, e_2)e_2, e_1) + g(R(e_1, e_3)e_3, e_1) = -2.$$

Similarly we have

$$S(e_2, e_2) = -2, \quad S(e_3, e_3) = 2 \quad \text{and} \quad S(e_i, e_j) = 0 \quad \text{for} \quad i \neq j.$$

Therefore,

$$r = S(e_1, e_1) + S(e_2, e_2) + S(e_3, e_3) = -2.$$

Hence we obtain that the scalar curvature is constant. Therefore from Theorem 4, it follows that M is a three-dimensional locally ϕ -conharmonically symmetric Sasakian manifold.

6. Conclusions

As a generalization of ϕ -symmetric Sasakian manifolds, ϕ -conharmonically symmetric Sasakian manifolds have been introduced in this paper. Conharmonic curvature tensor has some physical applications. Examples of three-dimensional locally ϕ -conharmonically symmetric Sasakian manifolds are given and prove that a ϕ -conharmonically symmetric Sasakian manifold is an Einstein manifold.

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