

Notes on some classes of 3-dimensional contact metric manifolds

J. E. Jin, J. H. Park and K. Sekigawa

Abstract. A review of the geometry of 3-dimensional contact metric manifolds shows that generalized Sasakian manifolds and η -Einstein manifolds are deeply interrelated. For example, it is known that a 3-dimensional Sasakian manifold is η -Einstein. In this paper, we discuss the relationships between several special classes of 3-dimensional contact metric manifolds which are generalizations of 3-dimensional Sasakian manifolds. We also provide examples illustrating our result in this paper.

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1 Introduction

It is well-known that any 3-dimensional compact oriented manifold admits a contact structure [21], and hence, it admits an associated contact metric structure. Therefore, it is natural to investigate 3-dimensional compact oriented manifolds from the contact metric view point. We shall give a brief review of contact metric manifolds focusing on the interrelationships between the generalizations of Sasakian manifolds and η -Einstein contact metric manifolds. It is well known that a Sasakian manifold is characterized as a contact metric manifold $M = (M, \phi, \xi, \eta, g)$ whose curvature tensor R satisfies

$$(1.1) \quad R(X, Y)\xi = \eta(Y)X - \eta(X)Y,$$

for any $X, Y \in \mathfrak{X}(M)$, where $\mathfrak{X}(M)$ denotes the Lie algebra of all smooth vector fields on M . As a generalization of the Sasakian manifold, Blair, Koufogiorgos and Papantoniou [2] introduced the notion of a contact metric manifold called a (κ, μ) -contact metric manifold satisfying the condition

$$(1.2) \quad R(X, Y)\xi = \kappa(\eta(Y)X - \eta(X)Y) + \mu(\eta(Y)hX - \eta(X)hY),$$

for any $X, Y \in \mathfrak{X}(M)$, where κ and μ are constants on M and $h = \frac{1}{2}\mathcal{L}_\xi\phi$ (here, \mathcal{L}_ξ is the Lie derivative in the direction of ξ). (κ, μ) -contact metric manifolds have attracted

by many authors [4, 5, 9, 10, 11, 18, 20]. (κ, μ) -contact metric manifolds include Sasakian manifolds ($\kappa = 1$ and $h = 0$), and also many examples of non-Sasakian (κ, μ) -contact metric manifolds have been provided. Koufogiorgos and Tsihlias [12] generalized the notion of a (κ, μ) -contact metric manifold by regarding the constants κ and μ in (1.2) to be smooth functions on M , called a *generalized (κ, μ) -contact metric manifold*. Further, the same authors [11] studied 3-dimensional generalized (κ, μ) -contact metric manifolds with $\xi\mu = 0$ (this condition means the function μ is constant along each integral curve of the characteristic vector field ξ) and showed that it is possible to construct two families of such manifolds in \mathbb{R}^3 , for any smooth function κ ($\kappa < 1$) of one variable. We shall introduce an example belonging to such families in §5, which illustrates Theorem B in the present paper. Koufogiorgos, Markellas and Papantoniou [10] introduced the notion of a (κ, μ, ν) -contact metric manifold which is a generalization of the generalized (κ, μ) -contact metric manifold, defined as a contact metric manifold $M = (M, \phi, \xi, \eta, g)$ satisfying

$$(1.3) \quad \begin{aligned} R(X, Y)\xi &= \kappa(\eta(Y)X - \eta(X)Y) + \mu(\eta(Y)hX - \eta(X)hY) \\ &\quad + \nu(\eta(Y)\phi hX - \eta(X)\phi hY). \end{aligned}$$

for any $X, Y \in \mathfrak{X}(M)$, where κ, μ, ν are smooth functions on M . In the same paper [10], they proved that a (κ, μ, ν) -contact metric manifold is necessarily a (κ, μ) -contact metric manifold if the dimension of M is greater than or equal to 5. They also proved that the condition (1.3) is invariant under the D -homothetic deformations, and further that, if $\dim M = 3$, then the condition (1.3) is equivalent to the following condition

$$(1.4) \quad Q = \left(\frac{r}{2} - \kappa\right)I + \left(-\frac{r}{2} + 3\kappa\right)\eta \otimes \xi + \mu h + \nu \phi h$$

holding on an open and dense subset of M , where Q is the Ricci operator and r is the scalar curvature of M ([10], Proposition 3.1). We note that $\kappa \leq 1$ on 3-dimensional (κ, μ, ν) -contact metric manifold (see(3.13)). A contact metric manifold $M = (M, \phi, \xi, \eta, g)$ is called η -Einstein if the Ricci operator Q takes the following form

$$(1.5) \quad Q = \alpha I + \beta \eta \otimes \xi,$$

where α and β are some smooth functions on M . From (1.3) and (1.4), taking account of (1.5), we may observe that the geometry of (κ, μ, ν) -contact metric manifolds and of generalized (κ, μ) -contact metric manifolds is deeply interrelated with the generalization of the η -Einstein contact metric manifold in the 3-dimensional case. On the other hand, a contact metric manifold $M = (M, \phi, \xi, \eta, g)$ is said to be H -contact if the characteristic vector field ξ is a harmonic vector field. We remark that (κ, μ, ν) -contact metric manifold is H -contact. Koufogiorgos, Markellas and Papantoniou [10] proved that a 3-dimensional H -contact manifold is a (κ, μ, ν) -contact metric manifold on an open and dense subset of M ([10], Theorem 1.1). The last two of the present authors worked on the H -contact unit tangent sphere bundles [6, 7, 14]. Concerning 3-dimensional (κ, μ, ν) -contact metric manifolds, the present authors previously proved the following theorem.

Theorem A [8] *Let $M = (M, \phi, \xi, \eta, g)$ be a 3-dimensional (κ, μ, ν) -contact metric manifold. If the functions μ and ν are constant on M , then M is either Sasakian*

or a non-Sasakian (κ, μ) -contact metric manifold with constant scalar curvature $r = 2\kappa - 2\mu$.

In this paper, we shall prove the following theorem.

Theorem B *Let $M = (M, \phi, \xi, \eta, g)$ be a 3-dimensional compact (κ, μ, ν) -contact metric manifold with $\xi\mu = \xi\nu = 0$ and let r be the scalar curvature. If either (the inequality) $r + \frac{\mu^2}{2} \geq 0$ or $r + \frac{\mu^2}{2} \leq 0$ holds everywhere on M , then M is a Sasakian manifold or a non-Sasakian (κ, μ) -contact metric manifold with $\kappa = \mu - \frac{\mu^2}{4}$ and $r = -\frac{\mu^2}{2}$.*

We here remark that the hypothesis “ $M = (M, \phi, \xi, \eta, g)$ is a 3-dimensional (κ, μ, ν) -contact metric manifold with $\xi\mu = \xi\nu = 0$ ” is preserved under any D -homothetic transformation [10] of the contact metric structure (ϕ, ξ, η, g) on M . Unless otherwise specified, the manifolds to be considered in this paper will be assumed to be connected.

2 Preliminaries

In this section, we present some basic facts about contact metric manifolds. We refer to [1] for more details. A $(2n+1)$ -dimensional smooth manifold M is called a *contact manifold* if it admits a global 1-form η such that $\eta \wedge (d\eta)^n \neq 0$ everywhere on M . We call η a *contact form* of M . It is well known that given a contact form η , there exists a unique vector field ξ , which is called the *characteristic vector field*, satisfying $\eta(\xi) = 1$ and $d\eta(\xi, X) = 0$ for any vector field X on M . A Riemannian metric g is said to be an *associated metric* to a contact form η if there exists a $(1, 1)$ -tensor field ϕ satisfying

$$(2.1) \quad \eta(X) = g(X, \xi), \quad d\eta(X, Y) = g(X, \phi Y), \quad \phi^2 X = -X + \eta(X)\xi,$$

where X and Y are vector fields on M . From (2.1), one can easily obtain

$$(2.2) \quad \phi\xi = 0, \quad \eta \circ \phi = 0, \quad g(\phi X, \phi Y) = g(X, Y) - \eta(X)\eta(Y).$$

The structure (ϕ, ξ, η, g) is called a *contact metric structure*, and a manifold M with a contact metric structure (ϕ, ξ, η, g) is said to be a *contact metric manifold* and is denoted by $M = (M, \phi, \xi, \eta, g)$. Let ∇ be the Levi-Civita connection and let R be the corresponding Riemann curvature tensor field given by $R(X, Y) = [\nabla_X, \nabla_Y] - \nabla_{[X, Y]}$ for all vector fields X, Y on M . We denote by S the Ricci tensor field of type $(0, 2)$, by Q the Ricci operator, and by r the scalar curvature. We define on M the operators h, l by setting

$$(2.3) \quad hX = \frac{1}{2}(\mathcal{L}_\xi \phi)X, \quad lX = R(X, \xi)\xi,$$

where \mathcal{L}_ξ is the Lie derivative in the direction of ξ . It is easily checked that h and l are symmetric operators and satisfy the following equalities

$$(2.4) \quad h\xi = 0, \quad l\xi = 0, \quad h\phi = -\phi h.$$

We also have the following formulas for a contact metric manifold:

$$(2.5) \quad \begin{aligned} \nabla_X \xi &= -\phi X - \phi hX, & (\text{and hence } \nabla_\xi \xi &= 0) \\ \nabla_\xi \phi &= 0, & Tr l &= g(Q\xi, \xi) = 2n - tr(h^2), \\ \phi l \phi - l &= 2(\phi^2 + h^2), & \nabla_\xi h &= \phi - \phi l - \phi h^2. \end{aligned}$$

On the other hand, a contact metric manifold for which ξ is a Killing vector field is called a *K-contact manifold*. It is well known that a contact metric manifold is *K-contact* if and only if $h = 0$. It is well known that Sasakian manifolds are necessarily *K-contact* but the converse is generally not true except in the 3-dimensional case ([1], pp.70 and pp.76). Here, we note that on any $(2n+1)(n > 1)$ -dimensional η -Einstein *K-contact* manifold, the functions α and β in the defining equation (1.5) are both constant. We may also note that any 3-dimensional Sasakian manifold is η -Einstein ((1.4), [17]) and $\alpha + \beta$ is constant [3]. Hence, it is natural to ask whether there exists a 3-dimensional Sasakian manifold with non-constant coefficient functions α and β as a η -Einstein or not. Concerning this question, to our knowledge, it seems that any explicit example of a 3-dimensional Sasakian manifold with non-constant coefficient functions α and β as an η -Einstein manifold has not yet appeared in any literature. In the last section, we shall provide an explicit example of such a 3-dimensional Sasakian manifold. Based on the above arguments, it seems worthwhile to discuss the coefficient functions in the equation (1.4) for a 3-dimensional (κ, μ, ν) -contact metric manifold, along with the generalizations of a 3-dimensional Sasakian manifold introduced in the §1.

3 Fundamental formulas

In this section, we shall prepare some fundamental formulas which we need in the proof of the Theorem B.

Let $M = (M, \phi, \xi, \eta, g)$ be a 3-dimensional contact metric manifold, and h, l be the $(1, 1)$ tensor fields defined by (2.3). First, we recall the following formula by [19]:

$$(3.1) \quad (\nabla_X \phi)Y = g(X + hX, Y)\xi - \eta(Y)(X + hX),$$

for any $X, Y \in \mathfrak{X}(M)$. Next, we recall that the curvature tensor R of a 3-dimensional Riemannian manifold satisfies the following identity

$$(3.2) \quad \begin{aligned} R(X, Y)Z &= g(Y, Z)QX - g(X, Z)QY - g(QX, Z)Y \\ &\quad + g(QY, Z)X - \frac{r}{2}(g(Y, Z)X - g(X, Z)Y), \end{aligned}$$

for any $X, Y, Z \in \mathfrak{X}(M)$. Now, let U be the open subset of M on which $h \neq 0$, and V be the open subset of points $m \in M$ such that $h = 0$ on a neighborhood of m . Then, we may easily check that $U \cup V$ is an open and dense subset of M . If U is not empty, for any $m \in U$, we may choose a local orthonormal frame field $\{\xi, e_1, e_2 = \phi e_1\}$ on a neighborhood of m in such a way that

$$(3.3) \quad h e_1 = \lambda e_1, \quad h e_2 = -\lambda e_2,$$

where λ is a smooth positive function on U . We may also note that, if V is not empty, then V becomes a Sasakian manifold (see §2).

Now, we assume that U is not empty. Then, by making use of (2.4), (2.5), (3.2) and (3.3), we have the following basic formulas on U :

$$(3.4) \quad \begin{aligned} \nabla_{\xi} e_1 &= -ae_2, & \nabla_{\xi} e_2 &= ae_1, & \nabla_{e_1} \xi &= -(\lambda + 1)e_2, & \nabla_{e_2} \xi &= -(\lambda - 1)e_1, \\ \nabla_{e_1} e_1 &= \frac{1}{2\lambda}(e_2\lambda + A)e_2, & \nabla_{e_1} e_2 &= -\frac{1}{2\lambda}(e_2\lambda + A)e_1 + (\lambda + 1)\xi, \\ \nabla_{e_2} e_2 &= \frac{1}{2\lambda}(e_1\lambda + B)e_1, & \nabla_{e_2} e_1 &= -\frac{1}{2\lambda}(e_1\lambda + B)e_2 + (\lambda - 1)\xi, \end{aligned}$$

and we have

$$(3.5) \quad [e_1, e_2] = -\frac{1}{2\lambda}(e_2\lambda + A)e_1 + \frac{1}{2\lambda}(e_1\lambda + B)e_2 + 2\xi,$$

where $A = S(\xi, e_1)$, $B = S(\xi, e_2)$ and a is a smooth function. Further, the Ricci operator Q [16] on U is given by

$$(3.6) \quad \begin{aligned} Q\xi &= 2(1 - \lambda^2)\xi + Ae_1 + Be_2, \\ Qe_1 &= A\xi + \left(\frac{r}{2} - 1 + \lambda^2 + 2a\lambda\right)e_1 + \xi(\lambda)e_2, \\ Qe_2 &= B\xi + \xi(\lambda)e_1 + \left(\frac{r}{2} - 1 + \lambda^2 - 2a\lambda\right)e_2. \end{aligned}$$

Thus, from (3.2) and (3.6), we get that the components of the curvature tensor are given by

$$(3.7) \quad \begin{aligned} R(e_1, e_2)e_1 &= \left(2 - \frac{r}{2} - 2\lambda^2\right)e_2 - B\xi, & R(e_1, e_2)e_2 &= \left(\frac{r}{2} - 2 + 2\lambda^2\right)e_1 + A\xi, \\ R(e_1, e_2)\xi &= Be_1 - Ae_2, & R(e_1, \xi)e_1 &= -Be_2 + (\lambda^2 - 1 - 2a\lambda)\xi, \\ R(e_1, \xi)e_2 &= Be_1 - \xi(\lambda)\xi, & R(e_1, \xi)\xi &= (2a\lambda + 1 - \lambda^2)e_1 + \xi(\lambda)e_2, \\ R(e_2, \xi)e_1 &= Ae_2 - \xi(\lambda)\xi, & R(e_2, \xi)e_2 &= Be_2 + (-1 + \lambda^2 + 2a\lambda)\xi, \\ R(e_2, \xi)\xi &= \xi(\lambda)e_1 + (1 - 2a\lambda - \lambda^2)e_2. \end{aligned}$$

We have noted that $Trl = 2(1 - \lambda^2)$ by (2.5). In the remaining section, we assume that M (under consideration) is a (κ, μ, ν) -contact metric manifold. Then, from (1.3), we have

$$(3.8) \quad R(e_1, e_2)\xi = 0, \quad R(e_1, \xi)\xi = (\kappa + \lambda\mu)e_1 + \lambda\nu e_2, \quad R(e_2, \xi)\xi = \lambda\nu e_1 + (\kappa - \lambda\mu)e_2.$$

Thus, comparing (3.7) and (3.8), we have

$$(3.9) \quad A = B = 0,$$

$$(3.10) \quad \xi\lambda = \lambda\nu,$$

$$(3.11) \quad 1 - \lambda^2 + 2a\lambda = \kappa + \lambda\mu, \quad 1 - \lambda^2 - 2a\lambda = \kappa - \lambda\mu,$$

Thus, from (1.4), (2.5), (3.6), (3.9), and (3.11), we have further

$$(3.12) \quad \mu = 2a,$$

$$(3.13) \quad \kappa = \frac{1}{2}S(\xi, \xi) = 1 - \frac{1}{2}Tr(h^2) = 1 - \lambda^2.$$

On the other hand, from (2.4) and (3.3), taking account of (3.4), (3.9), (3.10) and (3.12), we have

$$(3.14) \quad \begin{aligned} (\nabla_{e_1}\eta)(e_2) &= -(\lambda + 1), & (\nabla_{e_1}\eta)(\xi) &= 0, & (\nabla_{e_2}\eta)(e_1) &= -(\lambda - 1), \\ (\nabla_{e_2}\eta)(\xi) &= 0, & (\nabla_{\xi}\eta)(e_1) &= 0, & (\nabla_{\xi}\eta)(e_2) &= 0, \\ \\ (\nabla_{e_1}h)(e_2) &= -(e_1\lambda)e_2 + (e_2\lambda)e_1 - \lambda(\lambda + 1)\xi, \\ (\nabla_{e_2}h)(e_1) &= -(e_1\lambda)e_2 + (e_2\lambda)e_1 + \lambda(\lambda - 1)\xi, \\ (\nabla_{e_1}h)(\xi) &= -\lambda(\lambda + 1)e_2, & (\nabla_{e_2}h)(\xi) &= \lambda(\lambda - 1)e_1, \\ (\nabla_{\xi}h)(e_1) &= \lambda\nu e_1 - \lambda\mu e_2, & (\nabla_{\xi}h)(e_2) &= -\lambda\nu e_2 - \lambda\mu e_1, \\ (\nabla_{e_1}\phi h)(e_2) &= (e_1\lambda)e_1 + (e_2\lambda)e_2, & (\nabla_{e_1}\phi h)(\xi) &= \lambda(\lambda + 1)e_1, \\ (\nabla_{e_2}\phi h)(e_1) &= (e_1\lambda)e_1 + (e_2\lambda)e_2, & (\nabla_{e_2}\phi h)(\xi) &= \lambda(\lambda - 1)e_2, \\ (\nabla_{\xi}\phi h)e_1 &= \lambda\nu e_2 + \lambda\mu e_1, & (\nabla_{\xi}\phi h)e_2 &= \lambda\nu e_1 - \lambda\mu e_2. \end{aligned}$$

From (1.3), taking account of the second Bianchi identity, we get

$$(3.15) \quad \begin{aligned} & \mathfrak{S}_{X,Y,Z} R(X, Y)\nabla_Z\xi \\ &= \mathfrak{S}_{X,Y,Z} \{ (Z\kappa)(\eta(Y)X - \eta(X)Y) + \kappa((\nabla_Z\eta)(Y)X - (\nabla_Z\eta)(X)Y) \\ &+ (Z\mu)(\eta(Y)hX - \eta(X)hY) + \mu((\nabla_Z\eta)(Y)hX + \eta(Y)(\nabla_Zh)X \\ &- (\nabla_Z\eta)(X)hY - \eta(X)(\nabla_Zh)Y) + (Z\nu)(\eta(Y)\phi hX - \eta(X)\phi hY) \\ &+ \nu((\nabla_Z\eta)(Y)\phi hX + \eta(Y)(\nabla_Z\phi h)X - (\nabla_Z\eta)(X)\phi hY - \eta(X)(\nabla_Z\phi h)Y) \} \end{aligned}$$

for any $X, Y, Z \in \mathfrak{X}(M)$, where $\mathfrak{S}_{X,Y,Z}$ denotes the cycle sum with respect to the vector fields X, Y and Z . Setting $X = e_1, Y = e_2$ and $Z = \xi$ in (3.15), and taking account of (3.4), (3.7) and (3.14), we have

$$-2(\lambda^2 - 1 + \lambda^2\mu)\xi = 2(\kappa - \lambda^2\mu)\xi + (\lambda e_1\nu - \lambda e_2\mu - e_2\kappa)e_1 + (e_1\kappa - \lambda e_1\mu - \lambda e_2\nu)e_2,$$

and hence, we have

$$(3.16) \quad e_1\kappa = \lambda(e_1\mu + e_2\nu), \quad e_2\kappa = \lambda(e_1\nu - e_2\mu).$$

Thus, from (3.16), taking account of (3.13), we have also

$$(3.17) \quad e_1\lambda = -\frac{1}{2}(e_1\mu + e_2\nu), \quad e_2\lambda = \frac{1}{2}(e_2\mu - e_1\nu).$$

By the second Bianchi identity, we have further

$$(3.18) \quad \mathfrak{S}_{\xi, e_1, e_2} (\nabla_{\xi} R)(e_1, e_2)e_1 = 0,$$

Taking account of (3.4) and (3.7) with (3.9), (3.10), (3.12) and (3.13), we have

$$(3.19) \quad \begin{aligned} (\nabla_{\xi} R)(e_1, e_2)e_1 &= -\left(\frac{1}{2}\xi r + 4\lambda^2\nu\right)e_2, \\ (\nabla_{e_1} R)(e_2, \xi)e_1 &= -(e_1(\lambda\nu) + \mu e_2\lambda)\xi + \lambda(\lambda + 1)\nu e_2, \\ (\nabla_{e_2} R)(\xi, e_1)e_1 &= (e_2(\lambda\mu) - 2\lambda e_2\lambda + \nu e_1\lambda)\xi + \lambda(\lambda - 1)\nu e_2. \end{aligned}$$

Thus, from (3.18) and (3.19), we have

$$(3.20) \quad \xi r = -4\lambda^2\nu.$$

From (3.10) and (3.13), we have also

$$(3.21) \quad \xi\kappa = -2\lambda^2\nu.$$

Now, from (3.4), (3.9), (3.12) and (3.13), we obtain

$$(3.22) \quad \begin{aligned} R(e_1, e_2)e_1 &= \nabla_{e_1}(\nabla_{e_2}e_1) - \nabla_{e_2}(\nabla_{e_1}e_1) - \nabla_{[e_1, e_2]}e_1 \\ &= \left\{-\frac{1}{2}e_1(e_1 \log \lambda) - \frac{1}{2}e_2(e_2 \log \lambda) + \frac{1}{4}(e_2 \log \lambda)^2 + \frac{1}{4}(e_1 \log \lambda)^2 + \kappa + \mu\right\}e_2. \end{aligned}$$

On one hand, taking account of (2.5) and (3.4), we also obtain

$$(3.23) \quad \begin{aligned} &-\frac{1}{2}\Delta \log \lambda \\ &= -\frac{1}{2}\left\{e_1(e_1 \log \lambda) + e_2(e_2 \log \lambda) + \xi(\xi \log \lambda) - \frac{1}{2}(e_2 \log \lambda)^2 - \frac{1}{2}(e_1 \log \lambda)^2\right\}. \end{aligned}$$

Thus, from the first equality in (3.7), (3.22) and (3.23), we have

$$(3.24) \quad r = \Delta \log \lambda + 2\kappa - 2\mu - \xi\nu.$$

4 Proof of Theorem B

Let $M = (M, \phi, \xi, \eta, g)$ be a 3-dimensional compact (κ, μ, ν) -contact metric manifold with $\xi\mu = \xi\nu = 0$ on M . Now, we assume that the open subset U of M on which $h \neq 0$, is not empty. We set

$$(4.1) \quad \begin{aligned} F_{min} &= \{m \in M \mid \kappa \text{ takes into minimum at } m\}, \\ F_{max} &= \{m \in M \mid \kappa \text{ takes into maximum at } m\}. \end{aligned}$$

Then, we may easily check that F_{min} and F_{max} are both non-empty closed (and hence, compact) subsets of M such that $F_{min} \subset U$. And, we see that each integral curve of ξ is a geodesic in M . We denote by $\gamma(t) = \gamma(t; m)$ the integral curve of ξ through $m \in U$ with the arc-length parameter t . Then, from (3.10) and hypothesis $\xi\nu = 0$, we have

$$(4.2) \quad \lambda(t) \equiv \lambda(\gamma(t)) = \lambda(m)e^{\nu(m)t}.$$

for $|t| < \epsilon$, where ϵ is a certain positive real number. From (3.13), (4.2), we see that $\kappa(t) = \kappa(\gamma(t))$ is given by

$$(4.3) \quad \kappa(t) = 1 - \lambda(m)^2 e^{2\nu(m)t},$$

for $|t| < \epsilon$. Thus from (4.3), we see that, for each point $m \in U$, $\gamma(t) \in U$ for all $t \in \mathbb{R}$. Now, we suppose that there exists a point $m \in U$ with $\nu(m) > 0$. Then, from (4.3), we have

$$(4.4) \quad \lim_{t \rightarrow +\infty} \kappa(t) = -\infty.$$

Similarly, if there exists a point $m \in U$ with $\nu(m) < 0$. Then from (4.3), we have also

$$(4.5) \quad \lim_{t \rightarrow -\infty} \kappa(t) = -\infty.$$

Since M is compact, we see that $\kappa (\leq 1)$ must be bounded on M . But, from (4.4) and (4.5), this is a contradiction. Therefore, it follows that $\nu = 0$ on U . Since V is Sasakian, it follows immediately $\nu = 0$ on V . Since $U \cup V$ is an open and dense subset in M , we see that ν vanishes on M and hence, the (κ, μ, ν) -contact metric manifold M under consideration reduces to a generalized (κ, μ) -contact metric manifold with $\xi\mu = 0$. Since $\nu = 0$ on M , from (3.17), we have on U .

$$(4.6) \quad A_1 = -\frac{1}{2}B_1, \quad A_2 = \frac{1}{2}B_2,$$

where $A_1 = e_1\lambda$, $B_1 = e_1\mu$, $A_2 = e_2\lambda$, $B_2 = e_2\mu$. From (3.4) and (3.5), we have

$$(4.7) \quad [e_1, \xi] = \left(\frac{\mu}{2} - \lambda - 1\right) e_2, \quad [e_2, \xi] = -\left(\frac{\mu}{2} + \lambda - 1\right) e_1.$$

Since $\nu = 0$, from (3.10), we have also

$$(4.8) \quad \xi\lambda = 0.$$

Thus, from (4.7), taking account of (4.6) and (4.8), we obtain

$$(4.9) \quad \xi A_1 = \left(\lambda + 1 - \frac{\mu}{2}\right) A_2, \quad \xi A_2 = \left(\lambda - 1 + \frac{\mu}{2}\right) A_1.$$

Similarly, from (4.7), taking account of (4.6) and $\xi\mu = 0$, we obtain

$$(4.10) \quad \xi A_1 = -\left(\lambda + 1 - \frac{\mu}{2}\right) A_2, \quad \xi A_2 = -\left(\lambda - 1 + \frac{\mu}{2}\right) A_1.$$

Thus, from (4.9) and (4.10), we have

$$(4.11) \quad \left(\lambda + 1 - \frac{\mu}{2}\right) A_2 = 0.$$

$$(4.12) \quad \left(\lambda - 1 + \frac{\mu}{2}\right) A_1 = 0.$$

Lemma 4.1. $A_1 = 0$ or $A_2 = 0$ at each point of U .

Proof. We assume that $A_1 \neq 0$ and $A_2 \neq 0$ at some point $m \in U$. Then, from (4.11) and (4.12), it follows that $\lambda + 1 - \frac{\mu}{2} = 0$ and $\lambda - 1 + \frac{\mu}{2} = 0$ at the point m , and hence, $\lambda = 0$ at m . But, this is a contradiction. \square

Now, we define subsets F_1, F_2, G_1, G_2 and F of U by

$$\begin{aligned} G_1 &= \{m \in U \mid A_1 \neq 0 \text{ (i.e. } A_2 = 0) \text{ at } m\}, \\ G_2 &= \{m \in U \mid A_2 \neq 0 \text{ (i.e. } A_1 = 0) \text{ at } m\}, \\ F_1 &= \{m \in U \mid \lambda - 1 + \frac{\mu}{2} = 0 \text{ at } m\}, \\ F_2 &= \{m \in U \mid \lambda + 1 - \frac{\mu}{2} = 0 \text{ at } m\}, \\ F &= \{m \in U \mid A_1 = A_2 = 0 \text{ (i.e. } B_1 = B_2 = 0) \text{ at } m\}. \end{aligned}$$

Then, taking account of (4.11) and (4.12) and Lemma 4.1, we have the following relations.

$$(4.13) \quad \begin{aligned} G_1 \subset F_1, \quad G_2 \subset F_2, \quad F_1 \cap F_2 = \emptyset, \text{ and} \\ U = G_1 \cup G_2 \cup F = F_1 \cup F_2 \text{ (disjoint union)}. \end{aligned}$$

We have denoted by $F_{(i)}$ the interior of F in U . Then, taking account of (4.9), we may observe that, if $F_{(i)} \neq \emptyset$, then λ (and hence, κ) is constant on $F_{(i)}$. From (4.13), we see that $G_1 \cup G_2 \cup F_{(i)}$ is an open and dense subset in U . First, we assume that the inequality $r + \frac{\mu^2}{2} \geq 0$ holds on M . If $G_1 \neq \emptyset$, then from (3.24), taking account of (4.12), we have

$$(4.14) \quad \Delta \log \lambda = r - 2(1 - \lambda^2) - 4(\lambda - 1) = r + 2(\lambda - 1)^2 = r + \frac{\mu^2}{2} \geq 0$$

on G_1 . Similarly, if $G_2 \neq \emptyset$, then, from (3.24), taking account of (4.11), we have

$$(4.15) \quad \Delta \log \lambda = r - 2(1 - \lambda^2) + 4(\lambda + 1) = r + 2(\lambda + 1)^2 = r + \frac{\mu^2}{2} \geq 0$$

on G_2 . Therefore, we have the following inequality

$$(4.16) \quad \Delta \log \lambda \geq 0$$

on $G_1 \cup G_2$. By direct calculation, we get

$$(4.17) \quad \Delta \log \lambda = -\frac{1}{\lambda^2} |\text{grad} \lambda|^2 + \frac{1}{\lambda} \Delta \lambda$$

on $G_1 \cup G_2$. Further, since $\kappa = 1 - \lambda^2$ on U , we get also

$$(4.18) \quad \Delta\kappa = -2|\operatorname{grad}\lambda|^2 - 2\lambda \Delta\lambda$$

on $G_1 \cup G_2$. Thus, from (4.17) and (4.18), we have

$$(4.19) \quad \Delta\kappa = -4|\operatorname{grad}\lambda|^2 - 2\lambda^2 \Delta \log \lambda \leq 0$$

on $G_1 \cup G_2$. On the other hand, $\kappa = \text{const}$ on $F_{(i)}$. Since $G_1 \cup G_2 \cup F_{(i)}$ is an open and everywhere dense subset of U , from (4.19), we have the inequality $\Delta\kappa \leq 0$ on U . If $V \neq \emptyset$, V is Sasakian (and has $\kappa = 1$ on V), since $\kappa = 1$ on V , it is evident that $\Delta\kappa = 0$ holds on V . Since $U \cup V$ is open and everywhere dense in M , we see finally that

$$(4.20) \quad \Delta\kappa \leq 0$$

holds on M . On the other hand, the function κ takes its minimum on the non-empty subset F_{min} . Therefore, by Hopf's theorem, we see that κ is constant on M , and hence, μ is also constant on M . Next, we assume that the inequality $r + \frac{\mu^2}{2} \leq 0$ holds everywhere on M . Then, applying the similar arguments as in the previous case where $r + \frac{\mu^2}{2} \geq 0$, we have $\Delta\kappa \geq 0$ holds on M . Since the function κ takes its maximum on the non-empty subset F_{max} . Therefore, by Hopf's theorem, we see also that κ and μ are both constant on M .

As the result, we see that M is a non-Sasakian (κ, μ) -contact metric manifold with $\kappa = \mu - \frac{\mu^2}{4}$ and hence $r = -\frac{\mu^2}{2}$ by virtue of (3.24) if $U \neq \emptyset$. On the other hand, it is evident that M is Sasakian ($\kappa = 1$ and $\mu = \nu = 0$) if $U = \emptyset$. This completes the proof of Theorem B.

5 Examples

In this section, we shall provide an example of the 3-dimensional Sasakian manifold $M = (M, \phi, \xi, \eta, g)$ with non-constant coefficient functions α and β in the defining equation (1.5) of an η -Einstein manifold are both non-constant (see Example 1), and also an example of the 3-dimensional generalized (κ, μ) -contact metric manifold which illustrates as well as supports Theorem B (see Example 2). Example 1 below is a special case of the example introduced in Blair's book [1].

Example 1 Let $M = \mathbb{R}^3$ and set

$$(5.1) \quad \xi = 2 \frac{\partial}{\partial z}, \quad e_1 = 2 \frac{\partial}{\partial y}, \quad e_2 = 2 \left(\frac{\partial}{\partial x} - y^2 \frac{\partial}{\partial y} + y \frac{\partial}{\partial z} \right).$$

Let η be the 1-form dual to ξ , and define $(1, 1)$ -tensor field ϕ by $\phi\xi = 0$, $\phi e_1 = e_2$ and $\phi e_2 = -e_1$. Further, let g be the Riemannian metric defined by $g(\xi, \xi) = 1$, $g(\xi, e_i) = 0$ and $g(e_i, e_j) = \delta_{ij}$ for $1 \leq i, j \leq 2$. Then, by direct calculation, we may check that (M, ϕ, ξ, η, g) is a 3-dimensional Sasakian manifold and the Ricci transformation Q is given by

$$(5.2) \quad Q = -(2 + 24y^2)I + (4 + 24y^2)\eta \otimes \xi$$

on M . Therefore, from (5.2), we see that the 3-dimensional Sasakian manifold M provides an explicit example of the η -Einstein manifold with non-constant coefficient functions α and β in (1.5) which is mentioned in §2.

The following example which is constructed by Koufogiorgos and Tsihlias [11], which illustrates Theorem B.

Example 2 Let $M = \{(x, y, z) \in \mathbb{R}^3 | z > 0\}$ and set

$$(5.3) \quad \xi = \frac{\partial}{\partial x}, \quad e_1 = -2y \frac{\partial}{\partial x} + (2\sqrt{z}x - \frac{1}{4z}y) \frac{\partial}{\partial y} + \frac{\partial}{\partial z}, \quad e_2 = \frac{\partial}{\partial y}.$$

Let η be the 1-form dual to ξ , and define $(1, 1)$ -tensor field ϕ by $\phi\xi = 0$, $\phi e_1 = e_2$ and $\phi e_2 = -e_1$. Further, let g be the Riemannian metric defined by $g(\xi, \xi) = 1$, $g(\xi, e_i) = 0$ and $g(e_i, e_j) = \delta_{ij}$ for $1 \leq i, j \leq 2$. Then, by direct calculation, we may check that (M, ϕ, ξ, η, g) is a 3-dimensional generalized (κ, μ, ν) -contact metric manifold with $\kappa = 1 - z$, $\mu = 2(1 - \sqrt{z})$ (and $\nu = 0$) and $r + \frac{\mu^2}{2} = -\frac{5}{8z^2} < 0$ on M .

Thus, Example 2 shows that the compactness assumption in Theorem B plays an essential role.

It is well-known that a 3-dimensional Lie group G admits a discrete subgroup Γ such that the space of right cosets $\Gamma \backslash G$ is compact if and only if G is unimodular [13]. Let G be one of the following simply connected unimodular Lie groups: $\tilde{E}(2)$, $E(1, 1)$. Then, from the proof of the Theorem B and ([2, §4], [15]), we may check that $M = \Gamma \backslash G$ with a suitable discrete subgroup Γ of G , provides an example illustrating Theorem B for non-Sasakian case. **Acknowledgement.** This work was supported by the National Research Foundation of Korea (NRF) grant funded by the Korea government (MEST) (2011-0012987).

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Authors' addresses:

Jieun Jin

Department of Mathematics, Sungkyunkwan University Suwon 440-746, Korea.

E-mail: yosulbong@skku.edu

JeongHyeong Park

Department of Mathematics, Sungkyunkwan University Suwon 440-746, Korea, and

School of Mathematics Korea Institute for Advanced Study, Seoul 130-722, Korea.

E-mail: parkj@skku.edu

Kouei Sekigawa

Department of Mathematics, Niigata University, Niigata 950-2181, Japan.

E-mail: sekigawa@math.sc.niigata-u.ac.jp